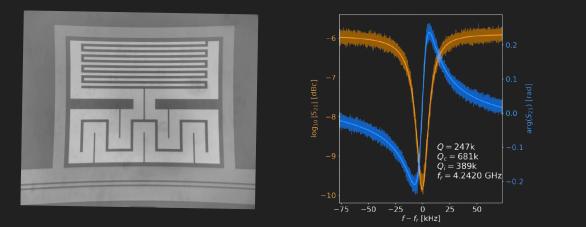
Kinetic inductance phonon-mediated (KIPM) detectors for low-mass dark matter

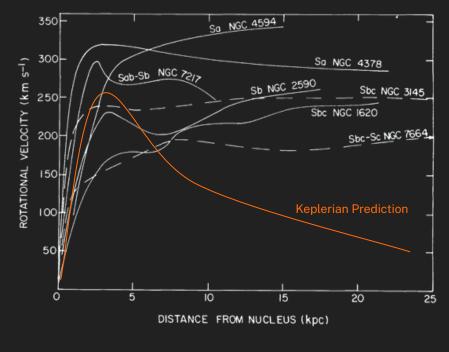




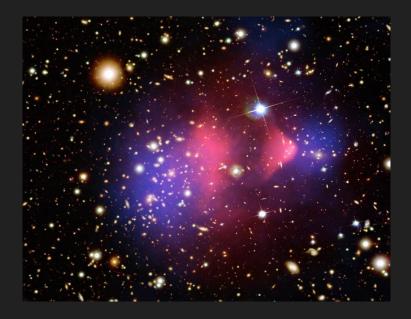
Brandon Sandoval, Advisor: Sunil Golwala

HEPCAT MEETING – NOVEMBER 1, 2024

Why does our universe point to dark matter?

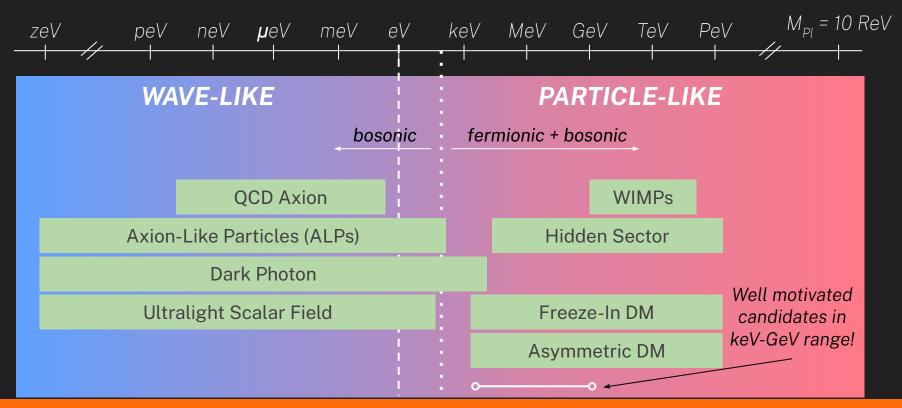


Rubin et al., 1978



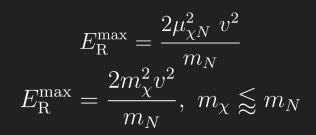
Cluster G.L.: Majority of mass in galaxies is "collisionless"

Dark matter mass spectrum



Direct detection kinematics

NUCLEAR RECOIL



ELECTRON RECOIL

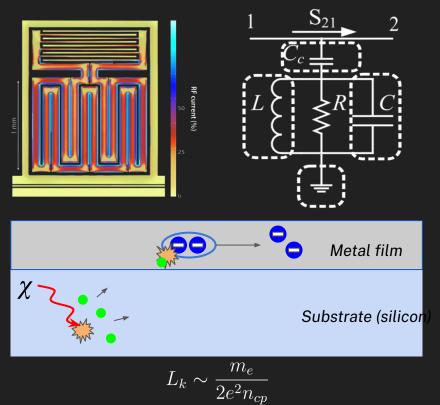
$$E_{\rm R} \sim m_e v_e v_{\chi}, \ m_{\chi} \gtrsim m_e$$
$$E_{\rm R} \sim \frac{1}{2} m_{\chi} v_{\chi}^2, \ m_{\chi} \lesssim m_e$$

 $v_{\chi} \sim 300 \text{ km/s}$

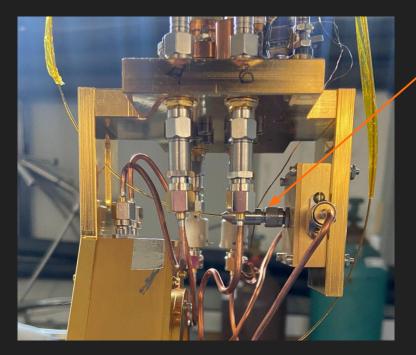
Dark matter with mass between 1 MeV to 1 GeV results in 0.1 meV to 100 eV deposited energy. Dominant energy excitation in this regime is phonons.

Kinetic Inductance Phonon-Mediated (KIPM) Detectors

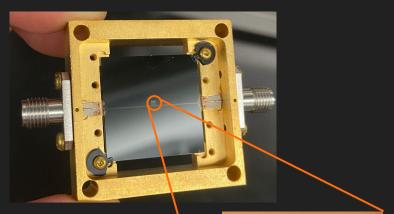
- Cooper pairs in superconducting film "lag" behind the field due to their inertia: kinetic inductance
- Energy deposition in substrate generates phonons, which travel to break cooper pairs, which increases the kinetic inductance
- By measuring the ratio of power sent into the LC resonator to the power on the output (S21), we can calculate the change in inductance, and hence the change in quasiparticle density and resulting energy deposited.



KID Deposited Energy Resolution Measurement

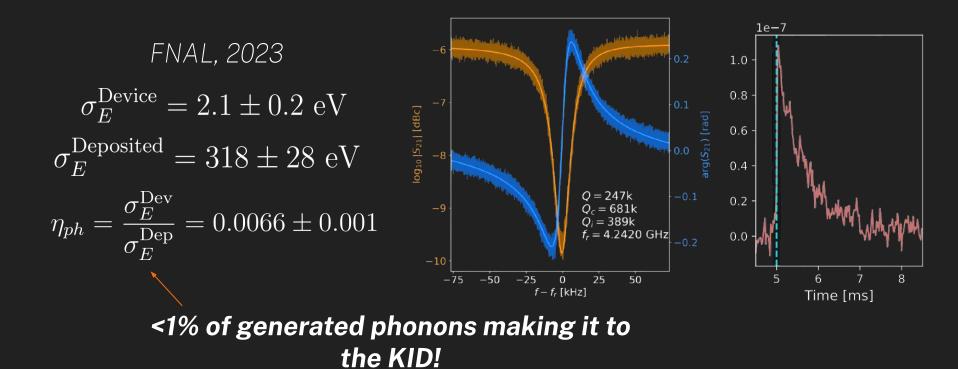


470 nm LED flashing the **back** of the silicon chip

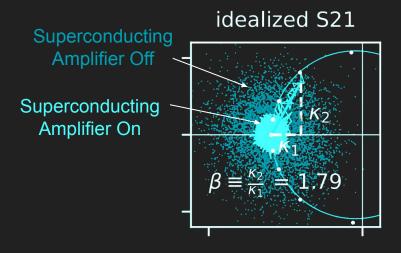




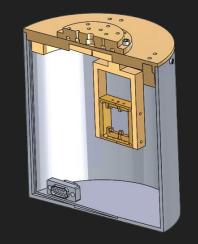
Energy resolution measurements: FNAL



KIPM current work: increase σ^{Device}



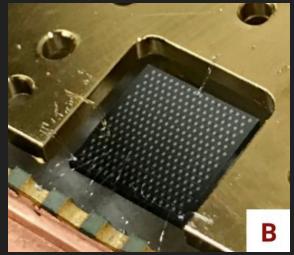
- Replacing semiconductor amplifier (HEMT) with superconducting amplifier
- Tests at JPL have shown a ~5x improvement in energy resolution



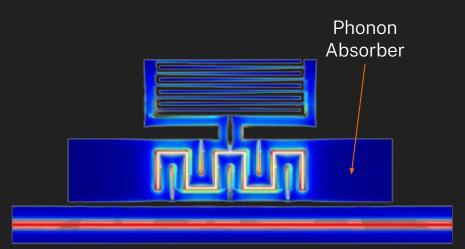
 Enhanced blackbody shielding – has been shown in partner facility in Fermilab to increase QP lifetime

KIPM current work: increasing η_{phonon}

Anthony-Petersen et al., 2022



• Designing new mounting scheme for wirebond-suspended chips



• Designing quasiparticle trapping KIDs with phonon absorbers

KIPM detector energy resolution projection

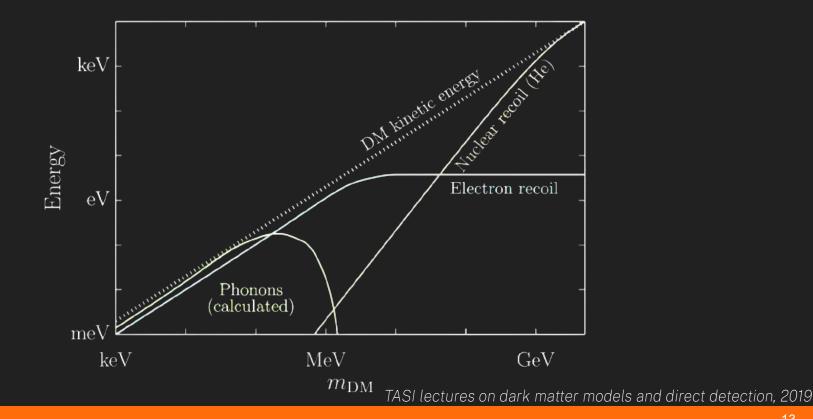


Thank You!

Questions?

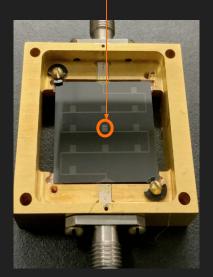
BACKUP SLIDES

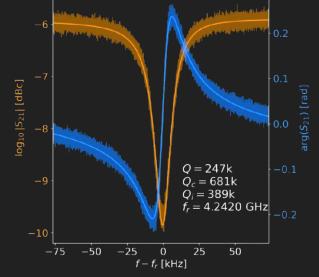
Energy deposition channels



Energy resolution measurement: FNAL

Phonon sensitive central KID

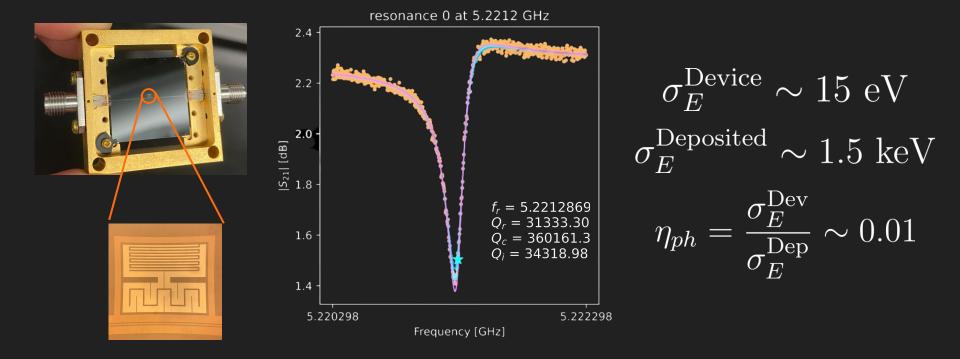


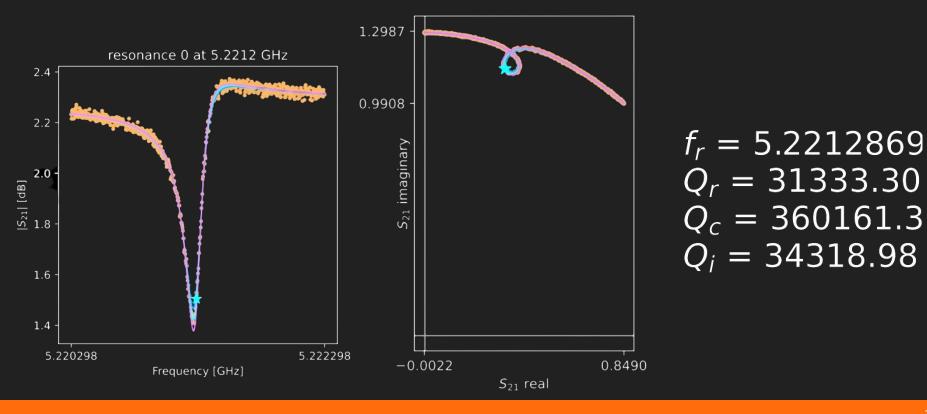


 $\sigma_E^{\text{Device}} = 2.1 \pm 0.2 \text{ eV}$ $\sigma_E^{\text{Deposited}} = 318 \pm 28 \text{ eV}$

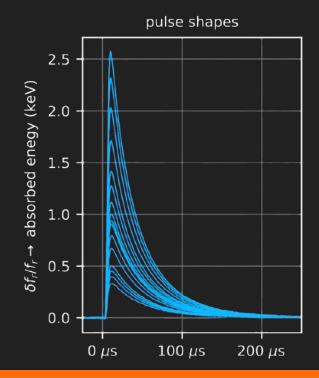
 $\eta_{ph} = \frac{\sigma_E^{\text{Dev}}}{\sigma_E^{\text{Dep}}} = 0.0066 \pm 0.001$

Energy resolution measurement: Caltech

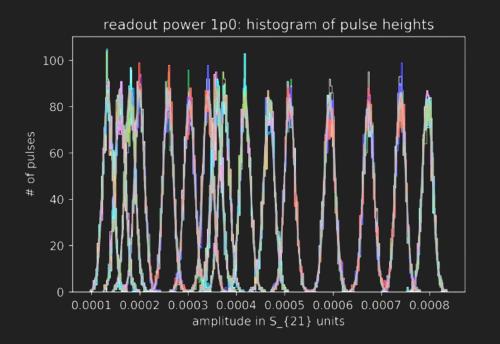


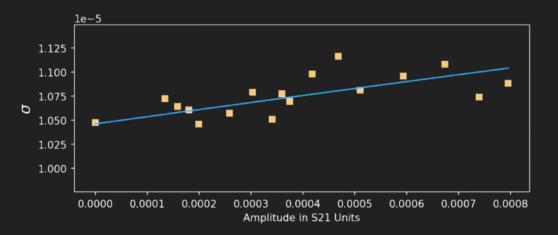


Average Pulses for Each LED Power



Histograms of Pulse Heights





$$\sigma^2 = \sigma_0^2 + r \cdot \mu$$

r is the responsivity per photon – exact expression depends on readout quadrature. Typical form is:

$$r = \alpha \frac{\kappa_i}{V} \frac{h\nu}{2\Delta} \eta_{ph}$$

$$\sigma_E^{\text{Deposited}} \sim 1.5 \text{ keV} \qquad \sigma_E^{\text{Device}} \sim 15 \text{ eV}$$
$$\eta_{ph} = \frac{\sigma_E^{\text{Dev}}}{\sigma_E^{\text{Dep}}} \sim 0.01$$

Kinetic Inductance Traveling Wave Parametric Amplifier (KI-TWPA)

$$\frac{\partial^2 I}{\partial z^2} - \frac{\partial}{\partial t} \left(LC \frac{\partial I}{\partial t} \right) = 0 \longrightarrow L(I) = L_0 \left(1 + \frac{I^2}{I_*^2} + \frac{I^4}{I_*^4} + \dots \right)$$

