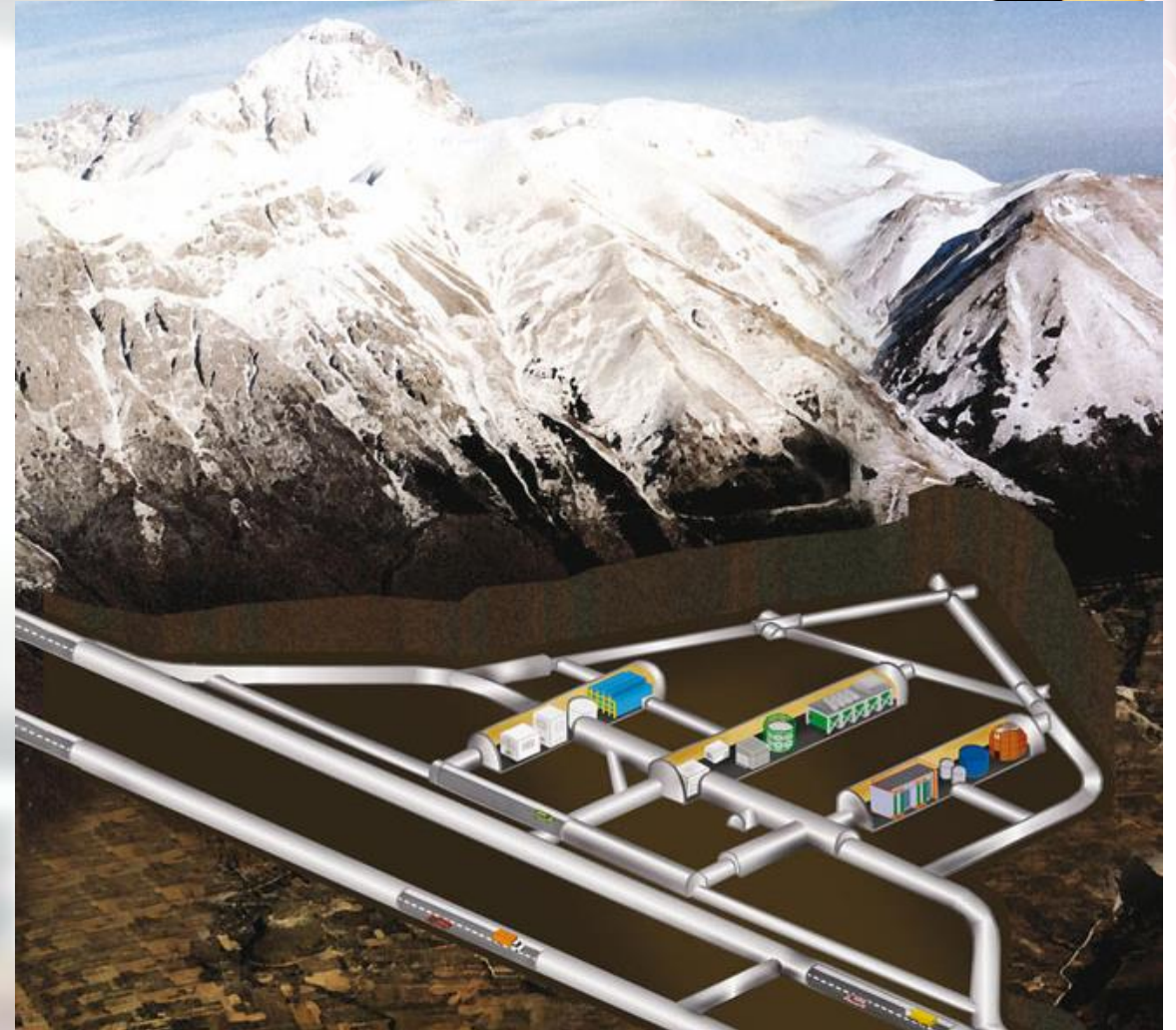


# Alec's HEP CAT Research Update: Improvements for Next-Generation Noble Liquid Detectors

HEPCAT 2024 Annual Meeting  
November 1, 2024, Alec Peck  
University of California, Riverside; Advisor: Dr. Shawn Westerdale  
& the DarkSide-20k Collaboration

## Gran Sasso National Laboratory (LNGS)

- Funded by the National Institute for Nuclear Physics (INFN) in Assergi, Italy
- Houses **DM direct detection** & neutrino physics experiments; XENON-nT, DAMA/LIBRA, BOREXINO, and...



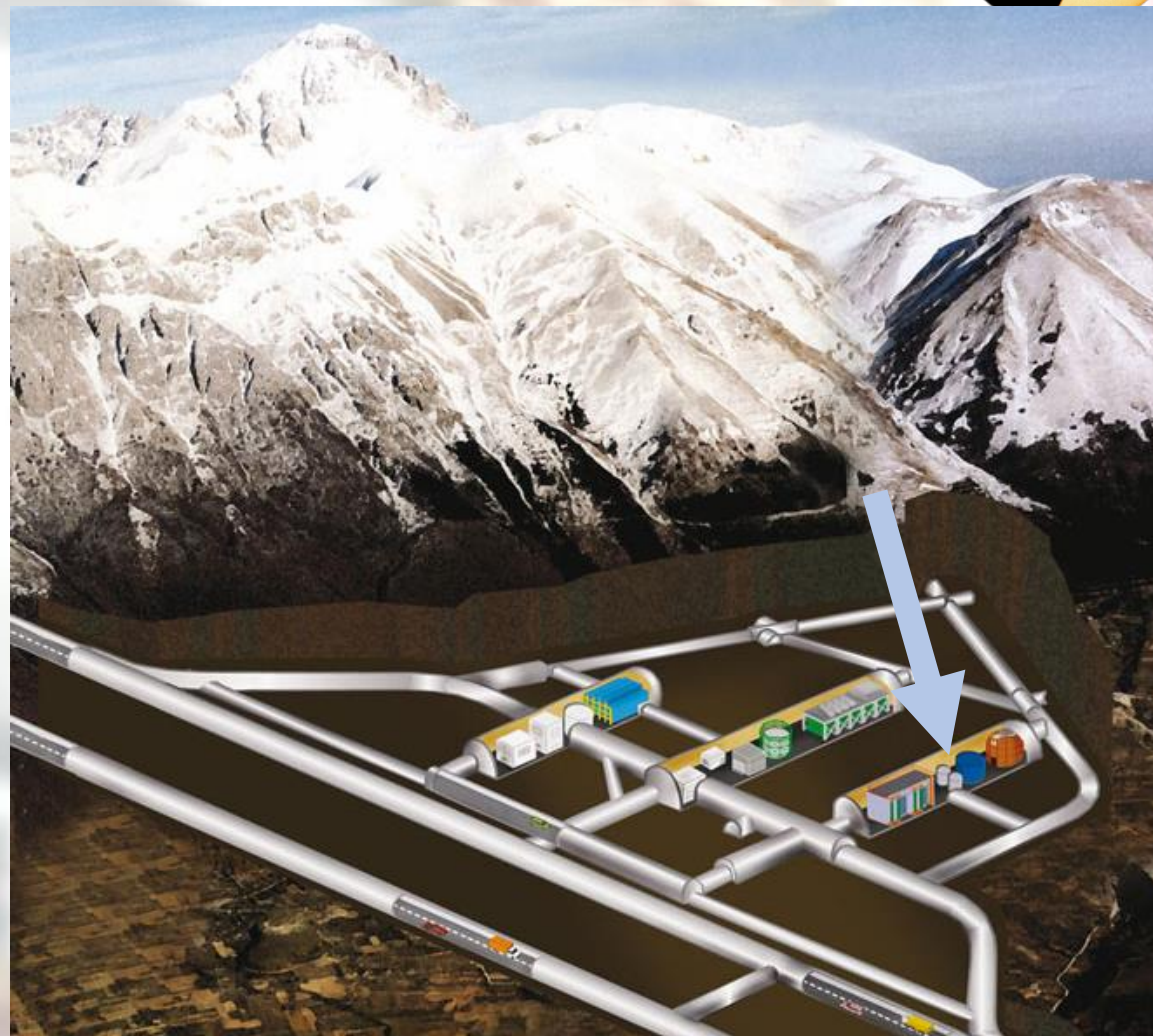
[https://indico.cern.ch/event/199223/contributions/378074/attachments/295975/413582/arneodo\\_lngs\\_red.pdf](https://indico.cern.ch/event/199223/contributions/378074/attachments/295975/413582/arneodo_lngs_red.pdf)

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## DarkSide-50 housed here

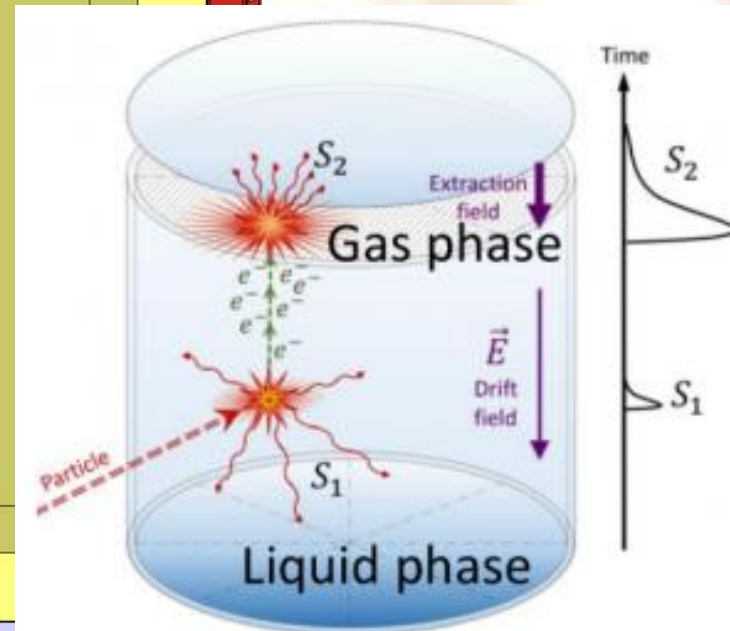
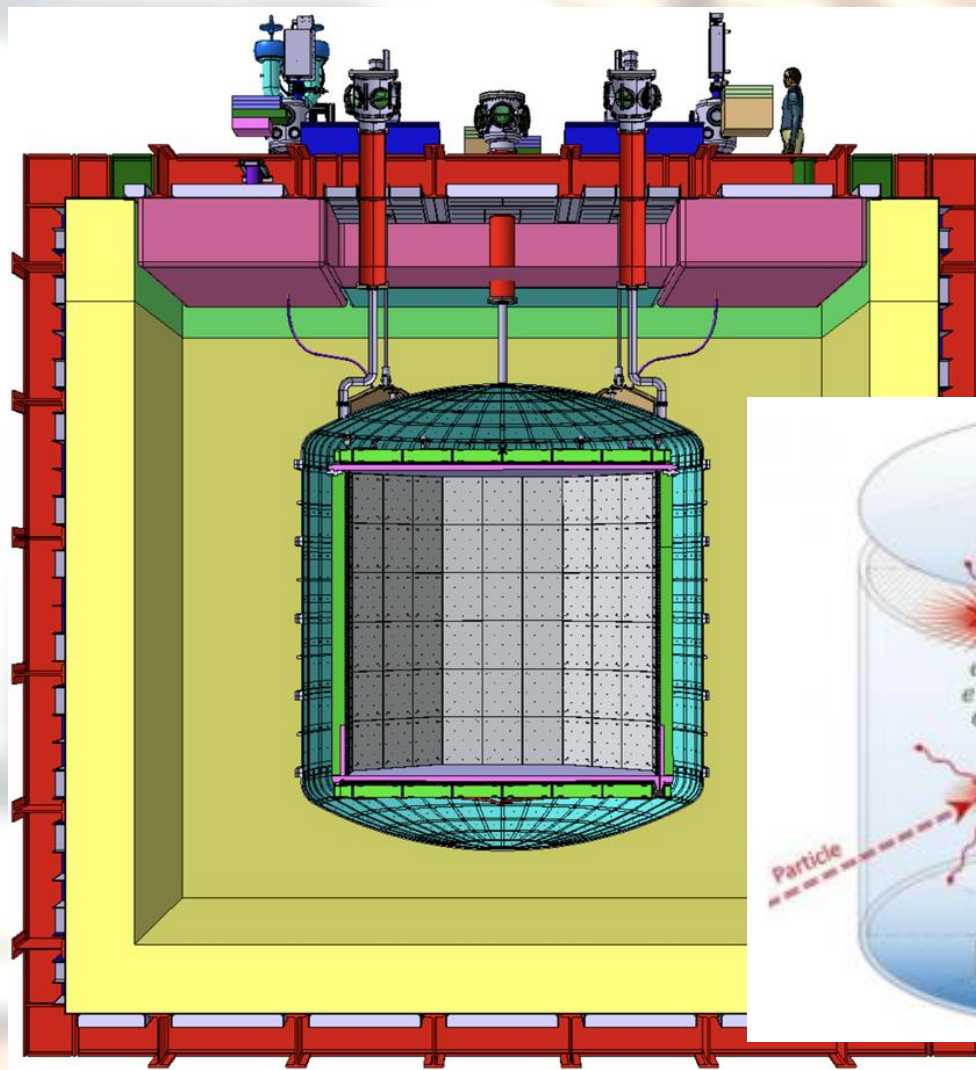
- Next generation: **DarkSide-20k**, **DarkSide-LM (Low-Mass)**
- DS20k under construction in Hall C for 2026



[https://indico.cern.ch/event/199223/contributions/378074/attachments/295975/413582/arneodo\\_lngs\\_red.pdf](https://indico.cern.ch/event/199223/contributions/378074/attachments/295975/413582/arneodo_lngs_red.pdf)

## Two-Phase TPC

- 50 tonnes of underground liquid argon with gas layer at the top
- Sensitivity to **WIMP-like interactions**
  - nucleon cross sections of  $\sim 10^{-24}$  barns ( $10^{-48}$  cm<sup>2</sup>)
  - Masses in 1 GeV to 1 TeV range
  - Neutrino physics, supernova detection
- **Scintillation light in liquid & ionization in gas gives energy and position information**



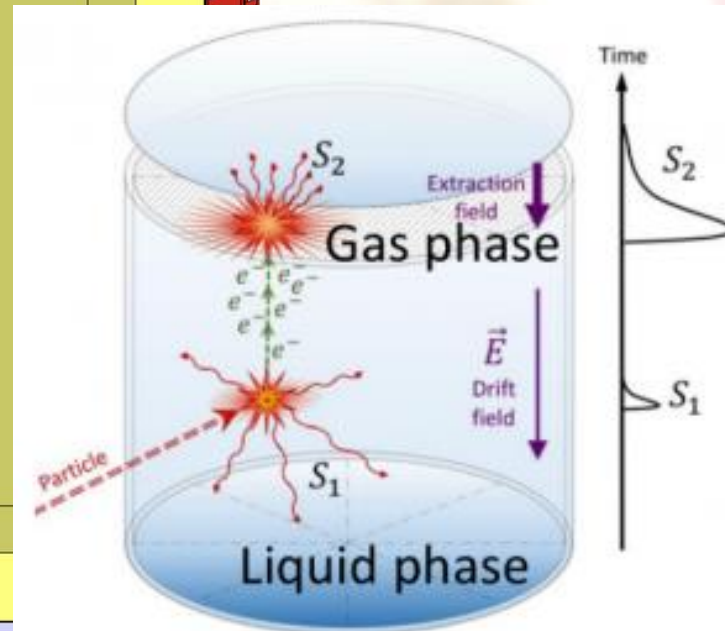
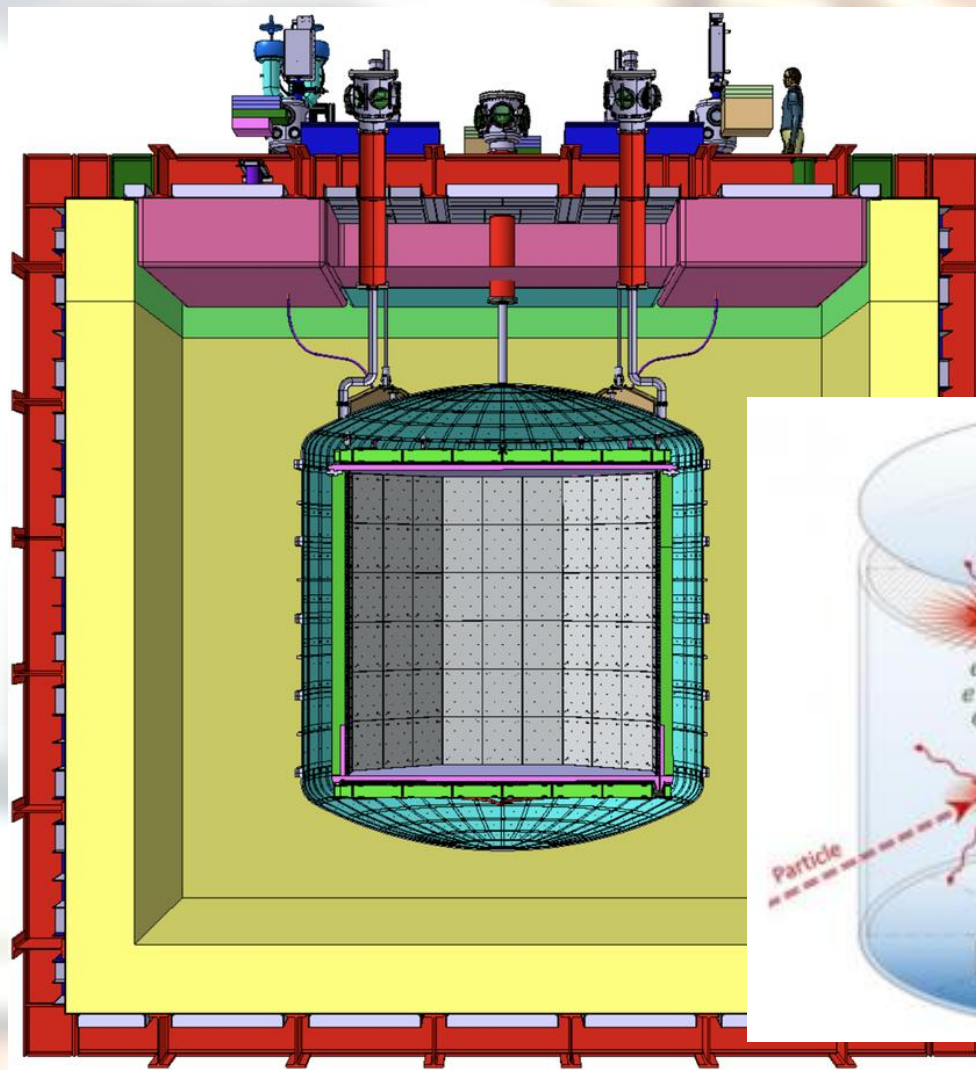
Note: image does not include HDPE neutron shielding or Outer Veto electronics  
<https://deap3600.ca/darkside-20k/>  
<https://www.lngs.infn.it/en/darkside>

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## Future Improvements

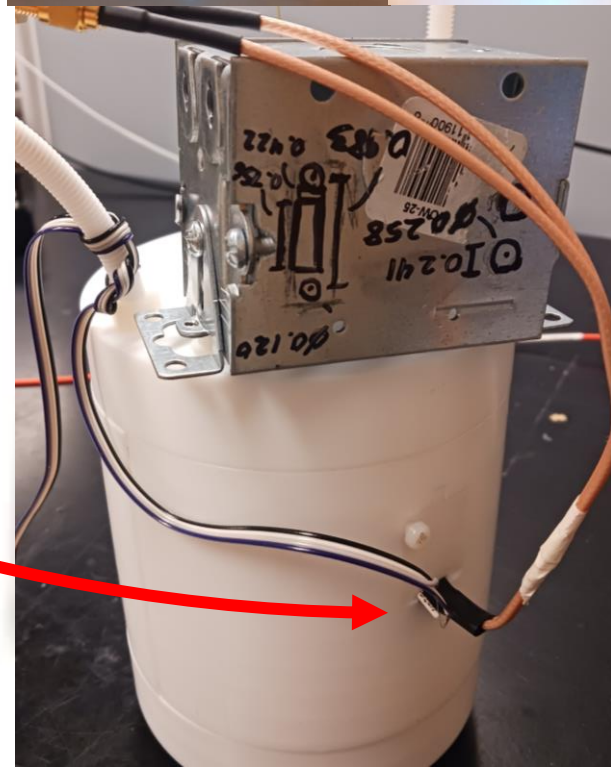
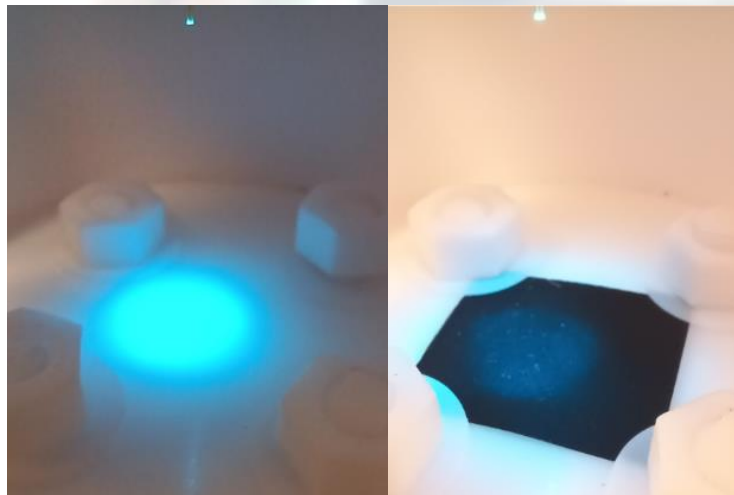
- Impact of **Xenon-doping** on GAR
- Test of **reflector material** optics
- Design and simulation of **outer veto**

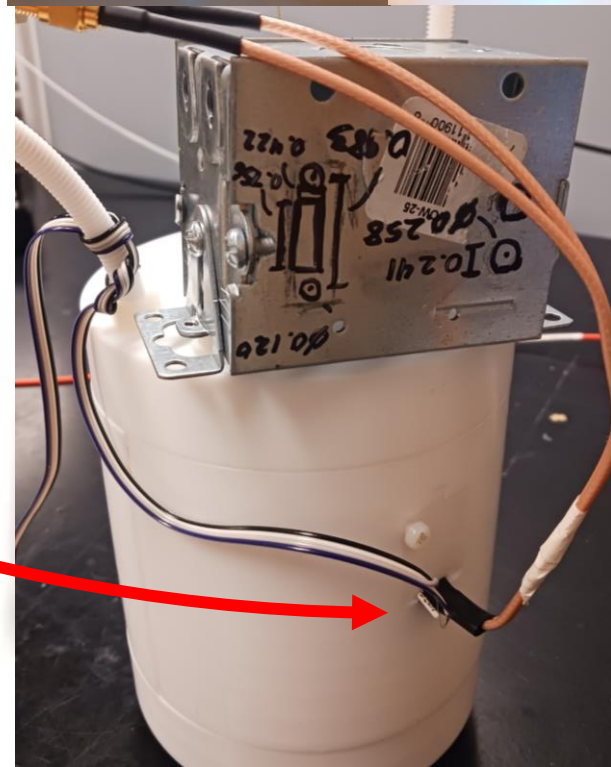
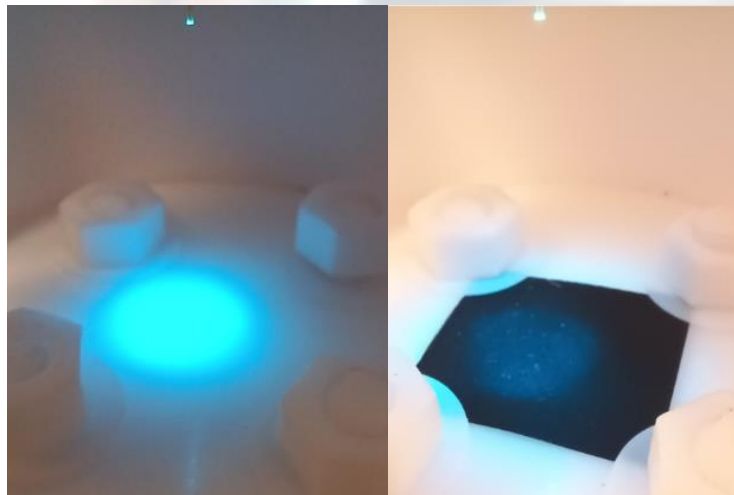


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## Reflectance Effects From LAr Exposure Characterization Tool

- PTFE test chamber “integrating sphere”
- UV/blue light delivered by optical fiber
  - Reflect off sample (black foil backing)
- Single SiPM unit captures **diffuse reflections**
  - Measure change in photon count for different reflectors





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## Cryogenic Conditions

- **Cold air** (100–150K): temp of reflectors in argon gas pockets
- **Cryogenic** (77–88K): temp of liquid argon
- Five broad wavelength LEDs to test reflectance spectra
  - Lumirror, Tyvek, ESR foil

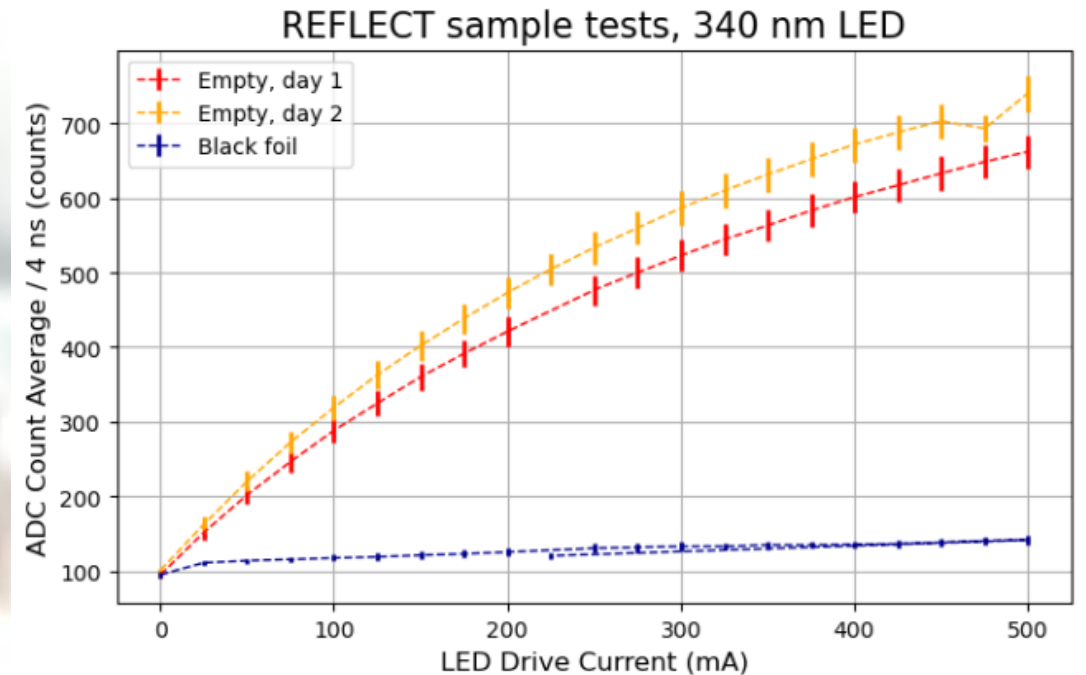
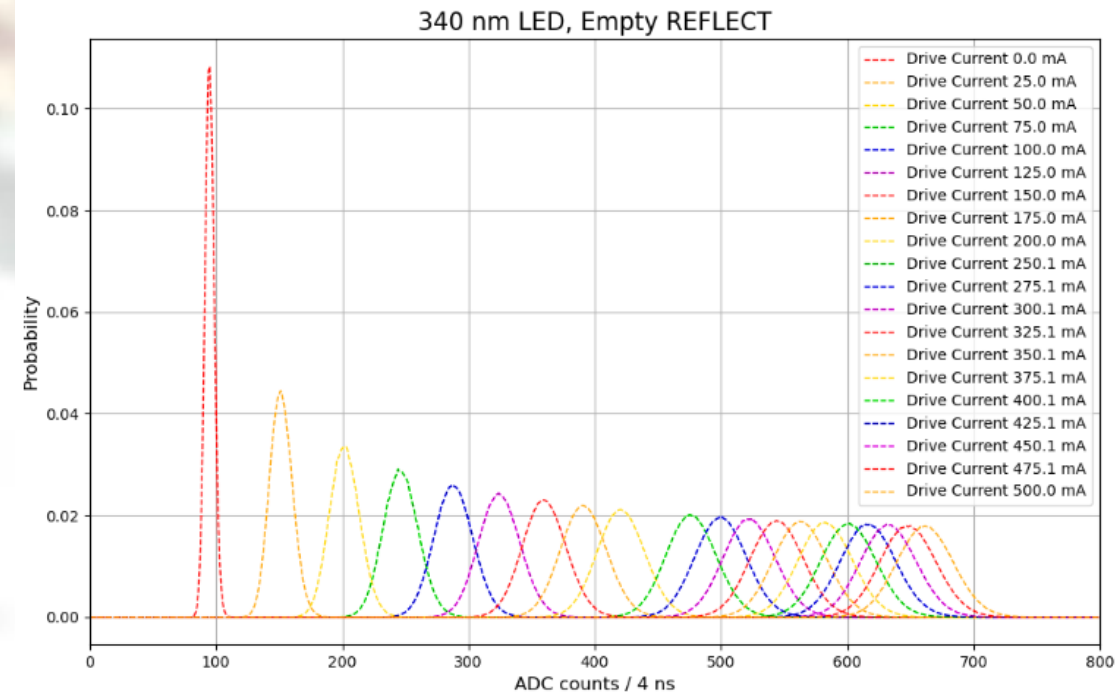


# Optics: REFLECT

- Signal strength v. optical power measured across linear SiPM dynamic range
- Calibrate empty REFLECT's response
  - Optimize optical design

## Next Steps

- Add reflector samples
- Low temperature (GAr and LAr temps)
- Reflectance spectra in 280–480 nm range

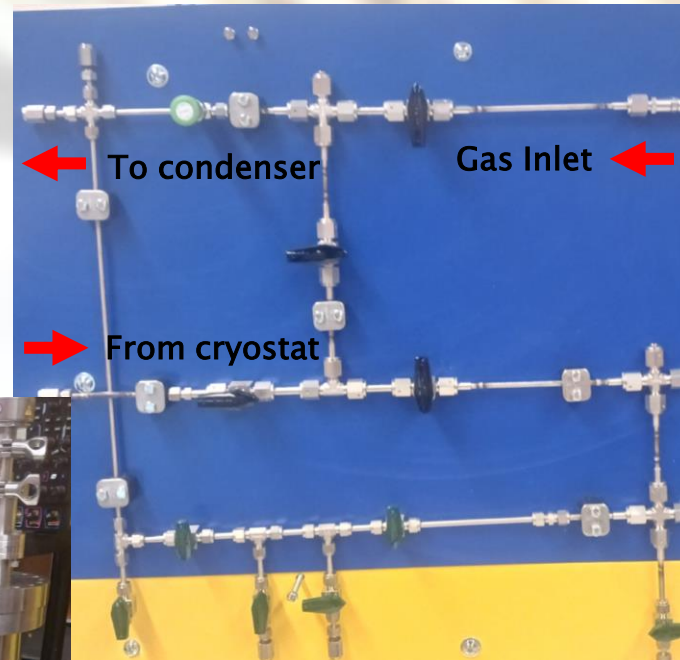




# Xenon Doping: LOADED

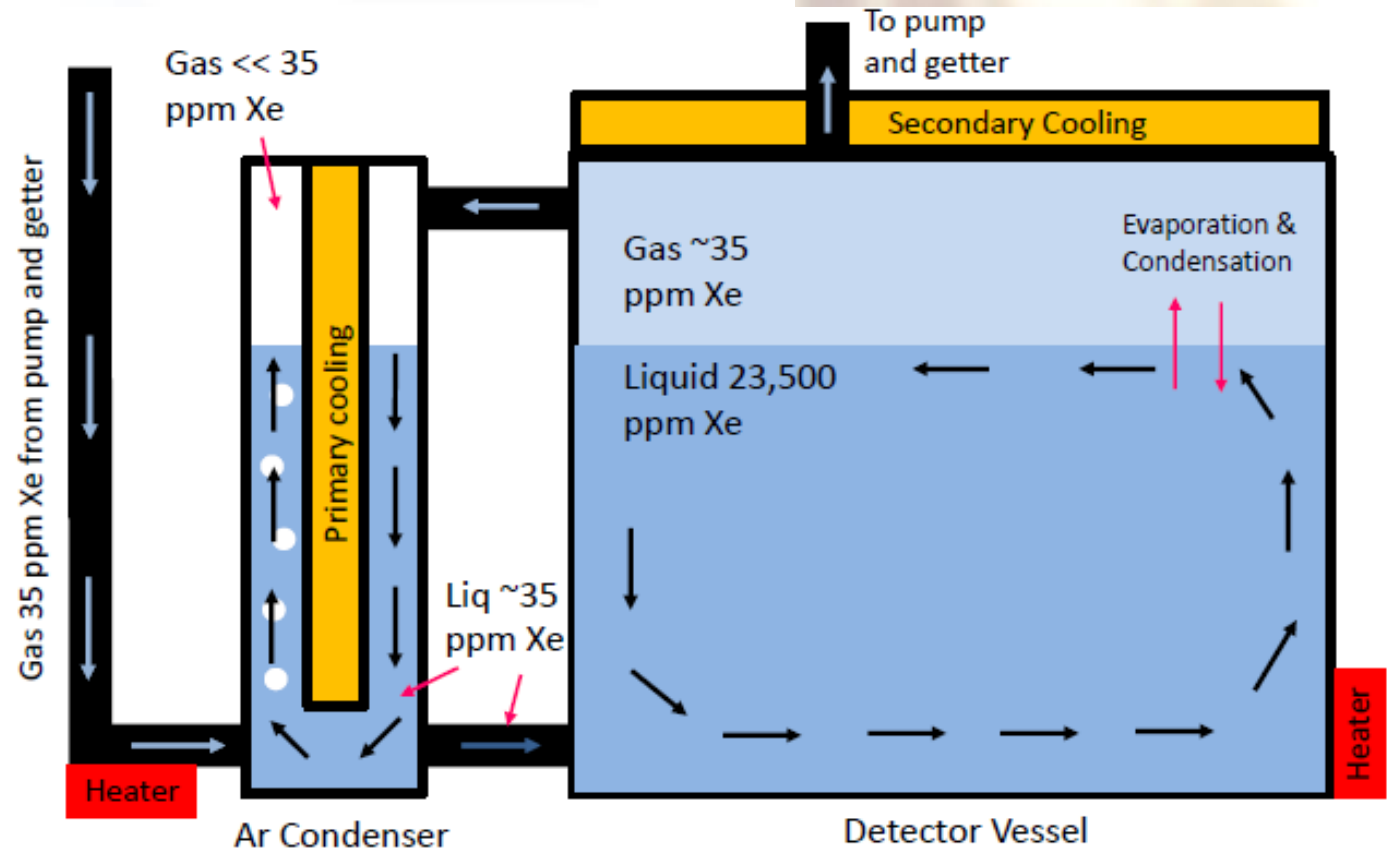
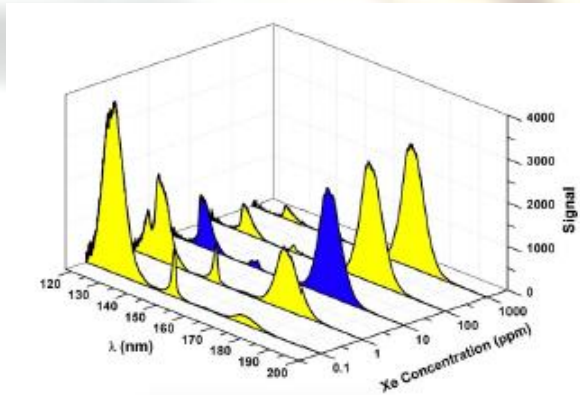
## Low-energy Optimization And Doping for Enhanced Detection

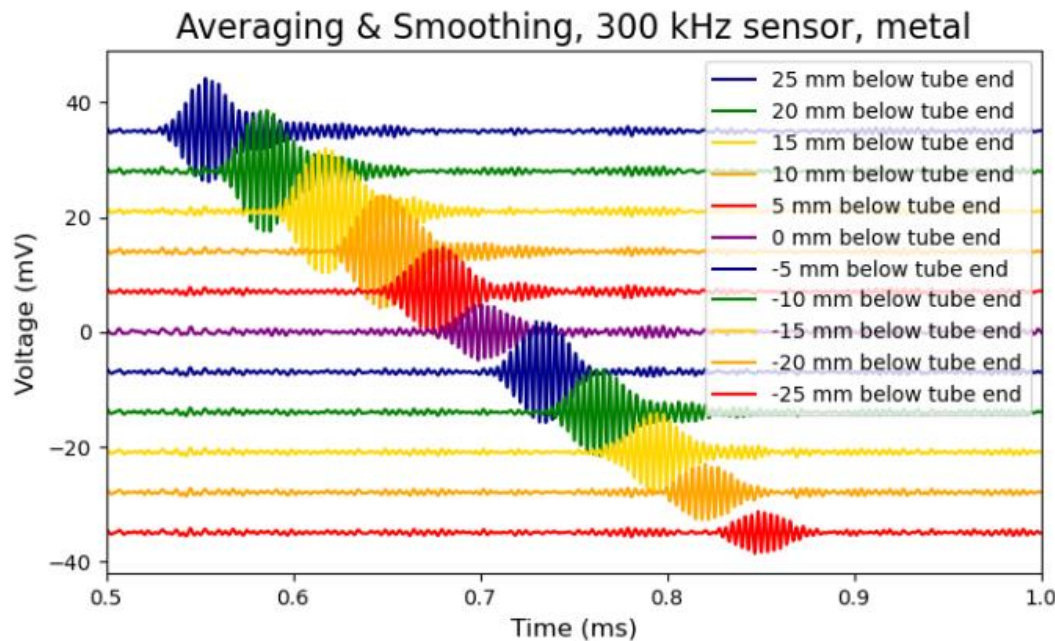
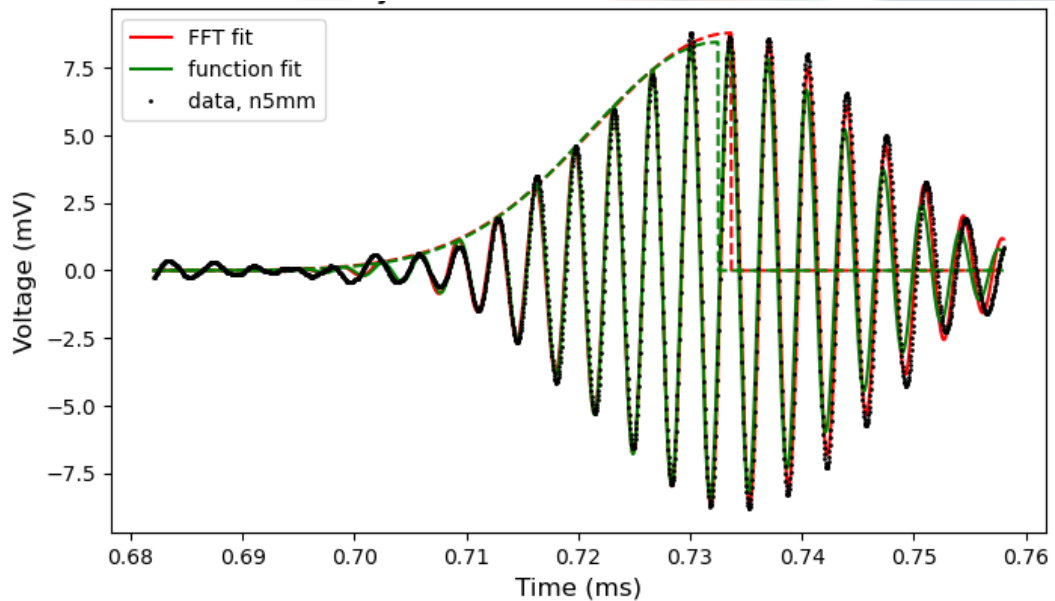
- Gas Panel mixes Ar/Xe and supplies to condenser
- Condenser liquifies Ar/Xe using  $N_2$ 
  - Xenon freezes before argon liquifies, must dissolve xenon
- Large cryostat to test xenon-doped argon bath



## Coherent Ionization Limit in Liquid Argon and Xenon

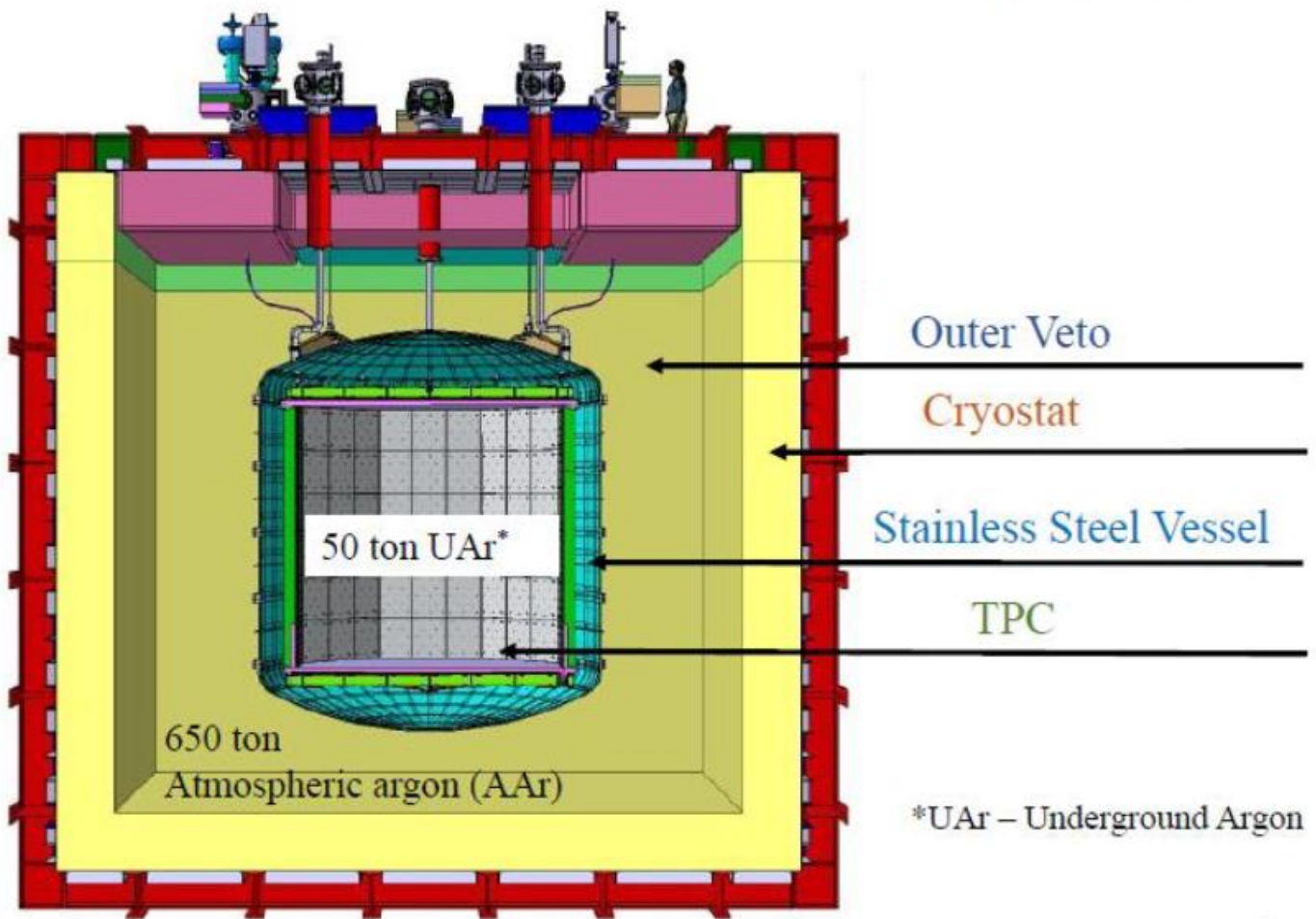
- Working with Dr. Jingke Xu's group at LLNL
  - LOADED based on previous work done here
  - Taking first Xe-doped LAr measurements
- Scalability testing
  - Thermodynamic stability for large detectors





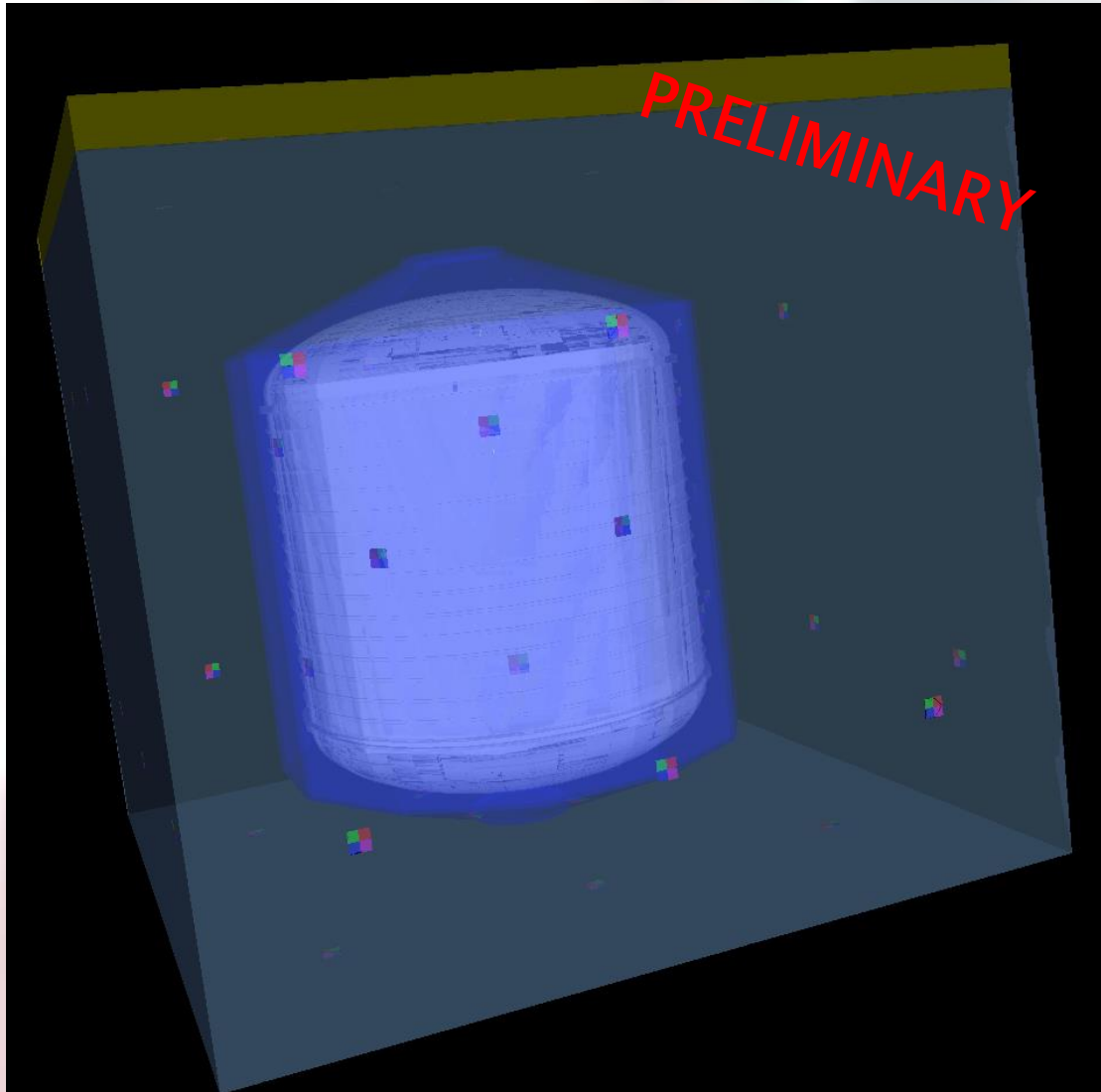
- My contribution: ultrasonic fill meter
  - Efficiency depends on width of gas pocket
  - Time of arrival of reflection pulse measures liquid level
  - Minimal heating
- Engineered pulse shape for sharp leading edge
- **Sub-mm precision** achieved in warm conditions

# DarkSide-20k: The Muon Veto



- 650 tons atmospheric argon , between the inner veto and Cryostat
  - Gas phase layer at top for stability
- **Passive shield** minimizes interactions
  - Cosmogenic neutrons generate muons and particle showers
- **Active scintillation detection**
  - Silicon Photomultiplier (SiPM) arrays give single-photon sensitivity
  - Tag cosmic ray events to veto for TPC analysis
  - Large volume gives potential for event detection
- Expected operation of DarkSide-20k by 2026

<https://www.lngs.infn.it/en/darkside>



- 32x16x24 (12,288) SiPM placed on interior of muon veto
- Argon scintillation light in UV (128 nm) is **wavelength shifted** to visible
- UV & visible light must be efficiently collected
  - Diffuse, **reflective interior** acts as large integrating sphere
  - Visible photons bounce between walls multiple times before detection

GEANT4 model of DarkSide-20k Cryostat. Liquid phase argon shown in light blue; gas phase argon shown in yellow; plastic shield shown in blue

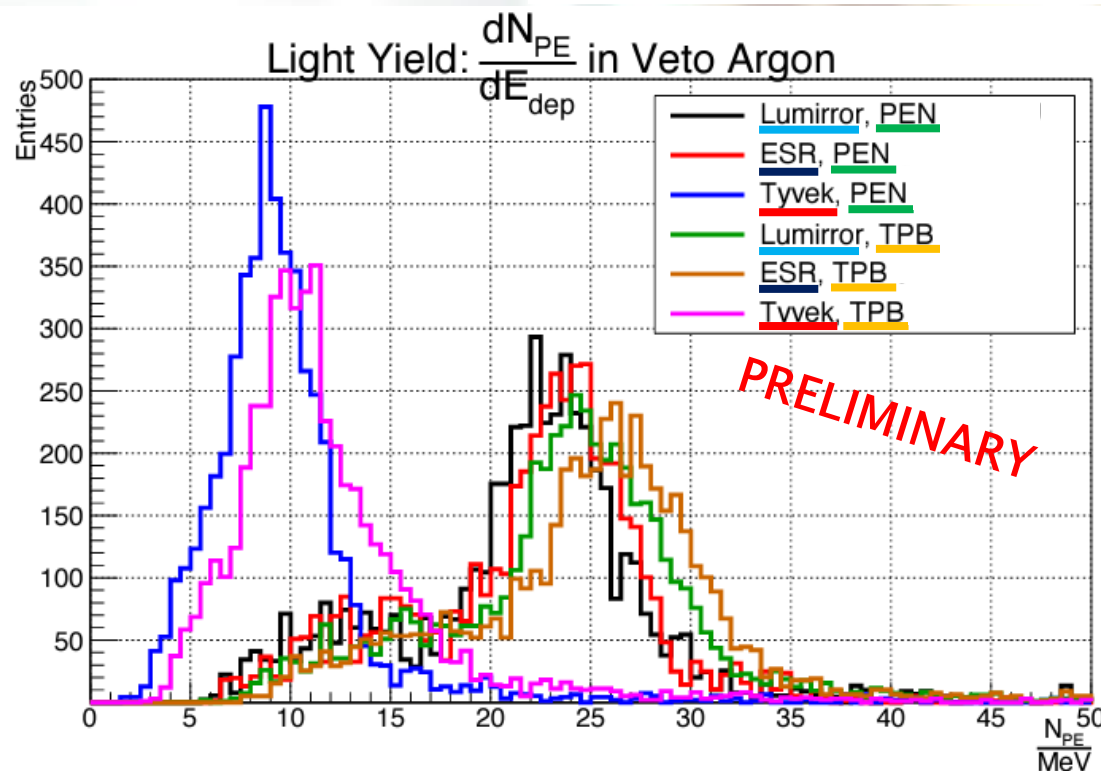
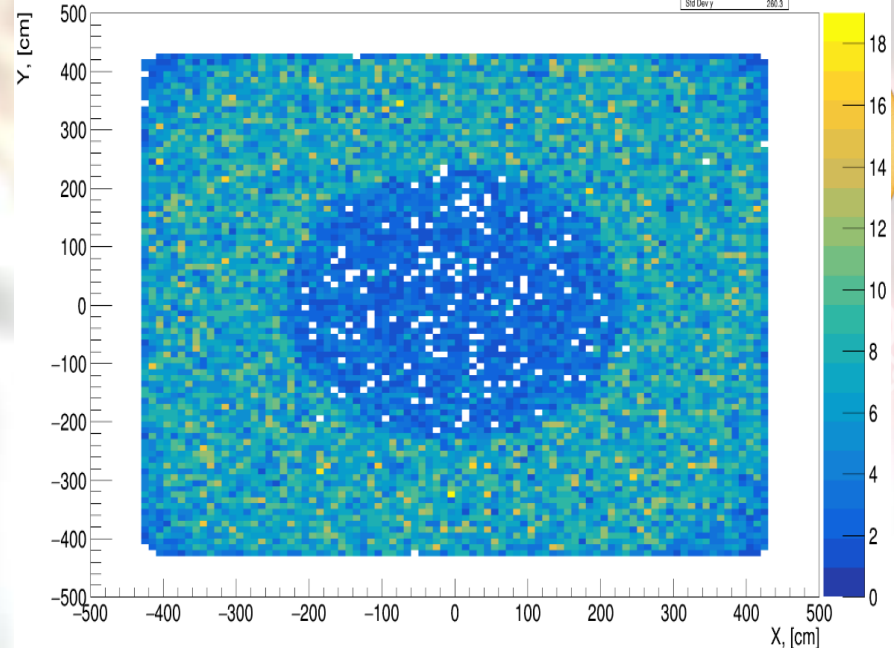


# Muon Veto Simulations

## Light Yield Optimization

- GEANT4 simulation of cosmic ray and background events
- Measure number of **detected** photons generated in all 32 PDU
- Light map specifies light yield per voxel for each individual sensor
- Confirming argon purity requirements

Work done by Taisiya Smirnova;  
University of California Riverside

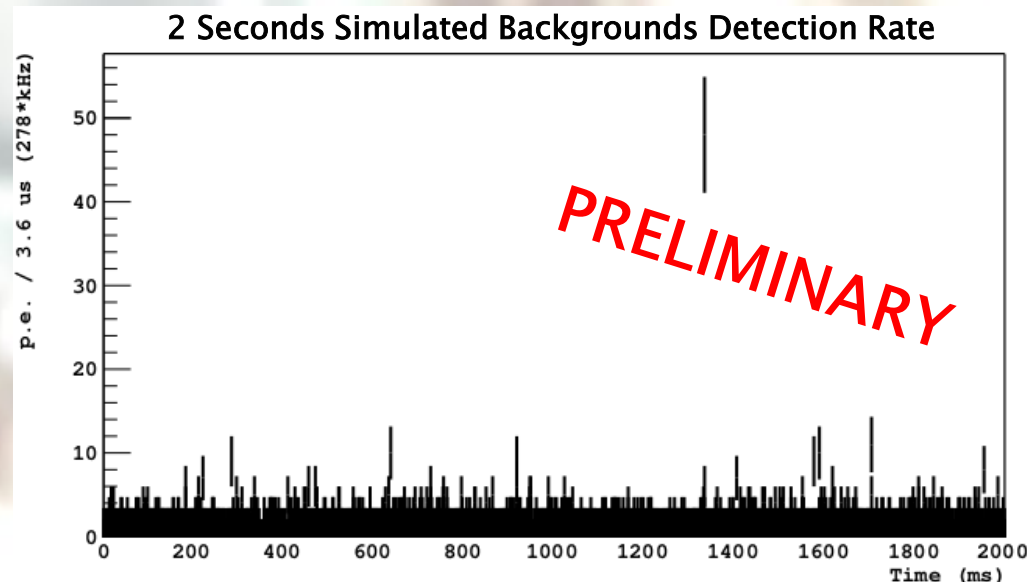
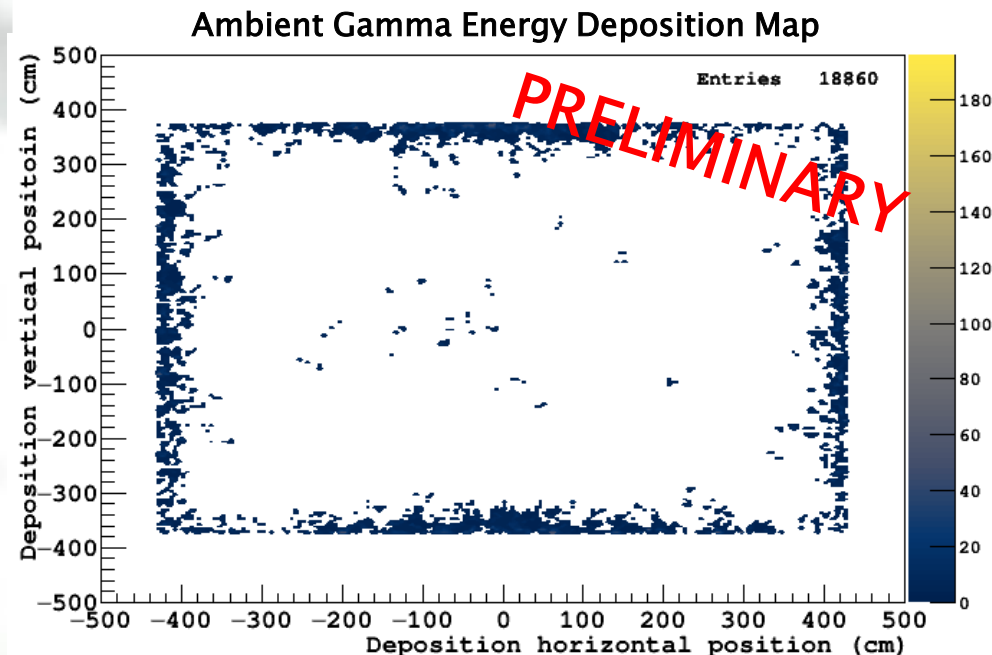


Work done by Ashlea Kemp;  
DarkSide Collaboration;  
University of Oxford

- Spatial and temporal distribution of radioactive backgrounds
  - Good passive shield: interactions do not reach inner veto
  - Low background rate: Will detect high energy muons above background

## Upcoming Work

- Estimation of data rates from active sensors
- Event detection analysis





# Thank you!



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